

**RESEARCHING  
REGOLITH**



**ON THE  
INTERNATIONAL SPACE STATION**

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00:00:00,350 --> 00:00:04,170

If you look at asteroids or if you look at the moon, you'll see this layer of

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00:00:04,170 --> 00:00:08,519

regolith covering the surface. So it's these fine particles, fine dust, that covers the

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00:00:08,519 --> 00:00:11,610

entire surface. So what's happening is the solar system is very dynamic and we

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00:00:11,610 --> 00:00:15,450

have asteroids and meteoroids moving around and impacting different bodies. So

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00:00:15,450 --> 00:00:18,750

that the surface is getting bombarded all the time, so you have this fine layer

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00:00:18,750 --> 00:00:20,440

of fragments on the surface.

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00:00:20,440 --> 00:00:22,050

So why would we want to study it, right? And why would

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00:00:22,050 --> 00:00:25,590

we want to study it on the space station? There's a lot of interest in one day

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00:00:25,590 --> 00:00:29,340

sending astronauts to asteroids, for instance, and we have already sent a lot

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00:00:29,340 --> 00:00:32,430

of robotic missions to asteroids a plan to send more. And so every time we visit

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00:00:32,430 --> 00:00:35,520

these asteroids we learn more and more  
and there's very diverse bodies of

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00:00:35,520 --> 00:00:37,350

asteroids. Every asteroid seems to be  
different.

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00:00:37,350 --> 00:00:41,080

And so we would need to understand how  
to interact with the asteroid surface.

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00:00:41,080 --> 00:00:44,879

So if you're going to anchor to the surface  
of the asteroid, for instance, we need to

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00:00:44,879 --> 00:00:48,600

understand how to interact with the  
regolith. So asteroids are pretty small

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00:00:48,600 --> 00:00:52,440

bodies and so they have actually  
microgravity levels of gravity, and so

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00:00:52,440 --> 00:00:55,100

this is actually what you find on the  
space station as well.

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00:00:55,120 --> 00:00:56,540

There's four experiments

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00:00:56,540 --> 00:00:59,520

and each one is a tube so  
it's a clear polycarbonate tube about

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00:00:59,520 --> 00:01:03,140

the size of a Pringles can, and so inside  
this tube we have the asteroid simulants,

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00:01:03,149 --> 00:01:06,210

and so this is what we're actually

studying and so the main piece of data

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00:01:06,210 --> 00:01:08,850

that we're getting is the movement of these particles. So we have a camera

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00:01:08,850 --> 00:01:12,119

looking at each of these tubes. Over the course of about a year that cassette one

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00:01:12,119 --> 00:01:15,150

will be up there we'll get this basically time-lapse imagery of the

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00:01:15,150 --> 00:01:18,479

particles moving, and so we're looking at the long-term effects of microgravity

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00:01:18,479 --> 00:01:20,840

and what's happened to the particles.

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00:01:20,840 --> 00:01:22,530

Also, something that we have on

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00:01:22,530 --> 00:01:26,790

asteroids is vacuum, right, so we want to expose our experiments to the vacuum of

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00:01:26,790 --> 00:01:31,400

space. So we actually have our facility hooked up essentially to a port to space.

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00:01:31,860 --> 00:01:35,400

We did cassette one but the idea is that Hermes will be open to other

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00:01:35,400 --> 00:01:38,820

investigators so cassette two, cassette 3, cassette 7, these will be other

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00:01:38,820 --> 00:01:41,579

researchers that will apply to use  
Hermes to do research on asteroid

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00:01:41,579 --> 00:01:44,970

regolith and granular material  
investigations, and we think that ISS is

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00:01:44,970 --> 00:01:48,299

a wonderful place to study asteroid  
regolith to study lunar regolith